LEO to GEO orbit design project report

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1 Problem

Inject spacecraft directly into circular parking orbit of altitude 100km. At burnout flight path angle is zero and velocity is V_{bo} .

Even though launch at equator (line of nodes), Orbit has an undesired 15^0 inclination in the geocentric frame and longitude of ascending node $\Omega = 20^0$. Need to correct this inclination so that you may rendezvous with a satellite in the GEO (35,860 km) on the equatorial plane with zero inclination.

This target satellite was 40^0 behind you at the time you entered your parking orbit of 100 km.

After you rendezvous with the first satellite, you need to transfer to a second satellite in GEO which was 10^0 ahead of you at time you entered your parking orbit.

Finally after making one complete orbit with this second satellite you need to take your final position in GEO which is another 5^0 ahead of the second satellite.

Design the above sequence. You may opt to minimize the time to complete or the fuel needed (ΔV). Specify your design criteria.

2 Assumptions

No drag while in LEO orbit. This allows the spacecraft to orbit as many times as needed to improve rendezvous conditions with the first target in GEO.

All impulses applied are assumed to have infinitely small time durations.

In addition, all assumptions used to derived Kepler equation apply as well.

Launch site latitude effect on ΔV are ignored. In practice, ΔV requirement need to be modified by small magnitude depending on the launch site location on the surface of the earth.

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3 Method

The geometry of the problem is illustrated in the figure below, which depicts the state of the system at t_0 . The time the space vehicle enters its parking orbit at burn out V_{bo} .

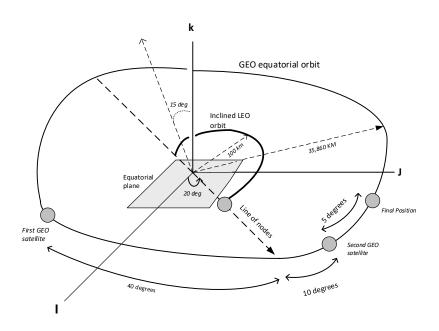


Figure 1: geometry of the problem

The analysis part will show an outline of the maneuvers to achieve the goal of the project.

The criteria for selecting the maneuvering sequence is:

The minimization of fuel

Which directly relates to the minimization of the ΔV .

4 Analysis

Now, we will show each phase of the maneuver, with different scenarios to achieve each phase. Before starting, there are common calculations that will be done now that will be shared by many scenarios, so that we do not have to re-calculate these each time.

4.1 Common calculations

For the Hohmann transfer from LEO to GEO:

Let V_p be the speed on the Hohmann orbit at the perigee point.

Let V_a be the speed on the Hohmann orbit at the apogee point.

Let 2a be the semi major axis for the Hohmann orbit.

Let T_h be the period of the Hohmann orbit.

For the LEO and GEO orbits:

Let V_{leo} be the speed on the LEO orbit.

Let V_{geo} be the speed on the GEO orbit.

Let r_{leo} be the radius of the LEO orbit.

Let r_{geo} be the radius of the GEO orbit.

Let T_{leo} be the period of the LEO orbit

let T_{geo} be the period of the GEO orbit

Then

$$\begin{split} r_{leo} &= 6378.145 + 100 = 6478.145 \text{ km} \\ r_{geo} &= 6378.145 + 35,860 = 42,238.145 \text{ km} \\ a &= \frac{r_{leo} + r_{geo}}{2} = \frac{6478.145 + 42238.145}{2} = 24358.145 \text{ km} \\ V_p &= \sqrt{\mu(\frac{2}{r_p} - \frac{1}{a})} \text{ but } r_p = r_{leo} \\ \text{so, } V_p &= \sqrt{3.986012 \times 10^5 \left(\frac{2}{6478.145} - \frac{1}{24358.145}\right)} = 10.3294 \text{ km/sec} = 37,185.84 \text{ km/hr} \\ V_a &= \sqrt{\mu(\frac{2}{r_a} - \frac{1}{a})} \text{ but } r_a = r_{geo} \\ \text{so, } V_a &= \sqrt{3.986012 \times 10^5 \left(\frac{2}{42238.145} - \frac{1}{24358.145}\right)} = 1.584 \text{ km/sec} = 5,702.4 \text{ km/hr} \\ V_{leo} &= \sqrt{\frac{\mu}{r_{leo}}} = \sqrt{\frac{3.986012 \times 10^5}{6478.145}} = 7.844 \text{ km/sec} = 28,238.82 \text{ km/hr} \\ V_{geo} &= \sqrt{\frac{\mu}{r_{geo}}} = \sqrt{\frac{3.986012 \times 10^5}{42238.145}} = 3.072 \text{ km/sec} = 11,059.089 \text{ km/hr} \\ 2a &= r_{leo} + r_{geo} = (6378.145 + 100) + (6378.145 + 35860) = 48,716.29 \text{ km}. \text{ Hence} \\ a &= 24,358.145 \text{ km}. \end{split}$$

$$T_{leo} = 2\pi \sqrt{\frac{r_{leo}^3}{\mu}} = 2\pi \sqrt{\frac{6478.145^3}{3.986012 \times 10^5}} = 5,189 \text{ sec} = 1 \text{ hr } 26 \text{ minutes and } 28 \text{ seconds.}$$

 $T_{geo} = 2\pi \sqrt{\frac{r_{geo}^3}{\mu}} = 2\pi \sqrt{\frac{42238.145^3}{3.986012 \times 10^5}} = 86,391 \text{ sec} = 23 \text{ hrs and 59 minutes and 50 seconds.}$

$$T_h = 2\pi \sqrt{\frac{a^3}{\mu}} = 2\pi \sqrt{\frac{24358.145^3}{3.986012 \times 10^5}} = 37,833 \text{ sec} = 10 \text{ hr } 30 \text{ min } 22 \text{ seconds}$$

4.2 Decide how to correct the LEO plane inclination

The problem is that we are given two non-coplanar circular orbits of different radices. A LEO orbit that is inclined at an angle to the plane of another, and larger circular GEO orbit.

We wish to transfer from the inclined $(i = 15^0)$ LEO orbit to the equatorial $(i = 0^0)$ GEO orbit.

We must correct the plane inclination to be able to transfer to the desired GEO orbit.

There are 3 possible ways to achieve this¹:

- 1. Correct all of the plane inclination before performing a Hohmann transfer from LEO to GEO. In other words, all of the inclination correction is made at the perigee of the Hohmann ellipse where the Hohmann transfer speed is largest. This will turn out to be the most fuel costly maneuver.
- 2. First perform a Hohmann transfer to transfer from the inclined LEO orbit to an inclined GEO orbit, and then apply all of the plane inclination correction at the apogee of the Hohmann elliptical orbit where the ellipse speed will be smallest. This is less costly in ΔV than the above sequence, and is a common maneuver.
- 3. Apply a small and partial plane inclination correction (say an angle α) at the perigee of the Hohmann orbit, then apply the remaining inclination correction (angle $\beta \alpha$) at the apogee. Notice that for $\alpha = 0$, this case becomes the same as case (2) above.

Now, we will analyze each case above in details and find the ΔV for each case and select the maneuver with the smallest ΔV .

¹Since the ratio of the radius of GEO orbit to radius of the LEO orbit is < 11.94, we do not need to consider using a bi-elliptical Hohmann transfer.

4.2.1 First scenario. All plane correction at perigee.

Move from the initial circular parking orbit (which has 15^0 degrees inclination) to a new circular orbit of the same radius but on the equatorial (0^0 degrees inclination). This requires one impulse to adjust the inclination. This impulse applied at the point where the parking orbit intersects the equator (line of nodes).

Next, and immediately, apply a coplanar Hohmann transfer (2 impulses) to transfer from the LEO orbit to the outer GEO orbit. (We Can combine the inclination correction velocity impulse vector with the first Hohmann velocity impulse vector using vector additions.)

This is illustrated in figure below.

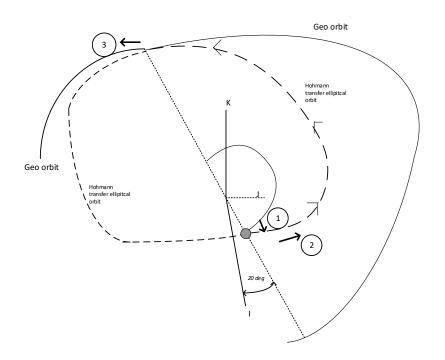


Figure 2: coplanar Hohmann transfer

To find impulse 1:

$$\Delta_1 V = 2V_{leo} \sin \frac{15^0}{2}$$

hence, $\Delta_1 V = 2(7.84412) \sin \frac{15^0}{2}$

$$\Delta_1 V = 2.048 km/sec = 7,371.81 km/hr$$

To find impulse 2:

$$\Delta_2 V = |V_l eo - V_p| = |7.844 - 10.3298| = 2.4858 km/sec = 8,948.88 km/hr$$

(speed up).

To find impulse 3:

$$\Delta_3 V = |V_g eo - V_a| = |3.072 - 1.584| = 1.488 km/sec = 5,356.8 km/hr$$

(speed up).

Hence, total impulses is found by summing the above

$$\Delta V = 2.048 + 2.4858 + 1.488 = 6.0218 km/sec = 21,678.48 km/hr$$

4.2.2 Second scenario. All plane correction at apogee.

Transfer from the initial LEO orbit (which has 15^{0} degrees inclination) to a GEO orbit at 35860 km altitude (which still has a 15^{0} degrees inclination). This is achieved using a normal Hohmann transfer (2 impulses). Next, perform an orbit plane inclination correction (one impulse) to move into the equatorial (0° inclination) GEO circular orbit on which the first target is currently orbiting.

This is illustrated in figure below.

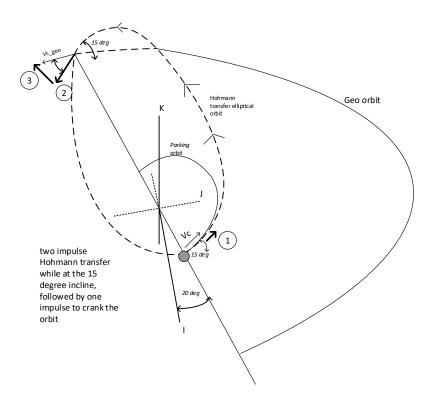


Figure 3: Transfer from the initial LEO orbit

To find impulse 1:

$$\Delta_1 V = V_{leo} - V_p = |7.844 - 10.3298| = 2.4858 \text{ km/sec} = 8,948.88 \text{ km/hr}$$

To find impulse 2:

$$\Delta_2 V = |V_{geo} - V_a| = |3.072 - 1.584| = 1.488 \text{ km/sec} = 5,356.8 \text{ km/hr}$$

(speed up). To find impulse 3

$$\Delta_3 V = 2V_{geo} \sin \frac{15^0}{2} = 0.80195 \text{ km/sec} = 2,887.031 \text{ km/hr}$$

Hence, total impulses is founding by summing the above total.

 $\Delta V = 2.4858 + 1.488 + 0.80195 = 4.77575 \text{ km/sec} = 17,192.7 \text{ km/hr}$

4.2.3 Third Scenario. Partial plane correction at perigee. Rest at apogee.

In this scenario, we will apply a partial orbit plane correction at the perigee and the remaining correction at the apogee. See figure below.

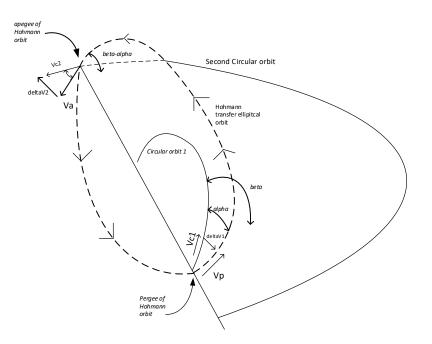


Figure 4: partial orbit plane correction

At the perigee of the Hohmann transfer, apply the law of the cosines to obtain

$$\Delta V_1 = \sqrt{V_{leo}^2 + V_p^2 - 2V_{leo}V_p\cos(\alpha)}$$
$$\Delta V_2 = \sqrt{V_{geo}^2 + V_a^2 - 2V_{geo}V_a\cos(\beta - \alpha)}$$
$$\Delta V_{total} = \Delta V_1 + \Delta V_2$$

To find the minimum ΔV_{total} for a given α , take α as the independent variable, and minimize ΔV_{total} as a function of α . Hence solve

$$\frac{\partial \Delta V_{total}}{\partial \alpha} = 0$$

Let

$$f(\alpha) = \Delta V_{total} = \sqrt{V_{leo}^2 + V_p^2 - 2V_{leo}V_p\cos(\alpha)} + \sqrt{V_{geo}^2 + V_a^2 - 2V_{geo}V_a\cos(\beta - \alpha)}$$

Hence for minimum

$$\frac{\partial f}{\partial \alpha} = \frac{2V_{leo}V_p\sin\left(\alpha\right)}{\sqrt{V_{leo}^2 + V_p^2 - 2V_{leo}V_p\cos\left(\alpha\right)}} - \frac{2V_{geo}V_a\sin\left(\beta - \alpha\right)}{\sqrt{V_{geo}^2 + V_a^2 - 2V_{geo}V_a\cos\left(\beta - \alpha\right)}} = 0$$

Hence

$$\frac{2V_{leo}V_p\sin(\alpha)\sqrt{V_{geo}^2 + V_a^2 - 2V_{geo}V_a\cos(\beta - \alpha)} - 2V_{geo}V_a\sin(\beta - \alpha)\sqrt{V_{leo}^2 + V_p^2 - 2V_{leo}V_p\cos(\alpha)}}{\sqrt{V_{leo}^2 + V_p^2 - 2V_{leo}V_p\cos(\alpha)}\sqrt{V_{geo}^2 + V_a^2 - 2V_{geo}V_a\cos(\beta - \alpha)}} = 0$$

Then

$$V_{leo}V_{p}\sin(\alpha)\sqrt{V_{geo}^{2} + V_{a}^{2} - 2V_{geo}V_{a}\cos(\beta - \alpha)} - V_{geo}V_{a}\sin(\beta - \alpha)\sqrt{V_{leo}^{2} + V_{p}^{2} - 2V_{leo}V_{p}\cos(\alpha)} = 0$$

Let

$$F(\alpha) = V_{leo}V_p\sin\left(\alpha\right)\sqrt{V_{geo}^2 + V_a^2 - 2V_{geo}V_a\cos\left(\beta - \alpha\right)} - V_{geo}V_a\sin\left(\beta - \alpha\right)\sqrt{V_{leo}^2 + V_p^2 - 2V_{leo}V_p\cos\left(\alpha\right)}$$

This is a non-linear equation in α . we solved for α using Newton root finding method. Hence we need to find $F'(\alpha)$ as follows

$$F'(\alpha) = V_{leo}V_p\cos(\alpha)\sqrt{V_{geo}^2 + V_a^2 - 2V_{geo}V_a\cos(\beta - \alpha)} + V_{leo}V_p\sin(\alpha)\left(\frac{-V_{geo}V_a\sin(\beta - \alpha)}{\sqrt{V_{geo}^2 + V_a^2 - 2V_{geo}V_a\cos(\beta - \alpha)}}\right)$$
$$-\left[V_{geo}V_a\cos(\beta - \alpha)\sqrt{V_{leo}^2 + V_p^2 - 2V_{leo}V_p\cos(\alpha)} + V_{geo}V_a\sin(\beta - \alpha)\frac{V_{leo}V_p\sin(\alpha)}{\sqrt{V_{leo}^2 + V_p^2 - 2V_{leo}V_p\cos(\alpha)}}\right]$$

To solve for α , I wrote a MATLAB function that uses Newton root finding method to find the root of $F(\alpha)$ for a given $r_{leo}, r_{geo}, \beta, \mu$ where β is the total angle (in degrees) of the inclination of the first circular orbit relative to the second circular orbit, and μ is the gravitational constant.

The function returns back the angle α for which ΔV_{total} is minimum.

For an initial guess for α , and since α is expected to be small compared to β , I selected $\alpha_0 = 0.1 \beta$.

Applying Newton iterative root finding:

$$\alpha_{i+1} = \alpha_i - \frac{F(\alpha_i)}{F'(\alpha_i)}$$

The tricky part in this problem was finding a good initial guess for the root (common problem with using Newton roots finding method).

I had to try different values for an initial guess before the root was converged to. For example, when I selected α_0 to be 50% of β , Newton method did not converge to the root. Selecting α_0 to be close to where one expects it to be (which is a small value compared to β) did work and a root was found.

For this design project, we are given that $\beta = 15^0$. Using this matlab function² I found that

$$\alpha = 1.28891^{\circ}$$

is the solution. Hence, this is the angle I will use for the correction to apply at the perigee.

So, at the perigee, Apply a correction of α angle, and at the apogee, apply a correction of 15 - 1.28891 = 13.711.

To find the impulse needed to correct inclination at the perigee and transfer from LEO to a Hohmann orbit with a correction of α

$$\Delta V_1 = \sqrt{V_{leo}^2 + V_p^2 - 2V_{leo}V_p\cos(\alpha)} = \sqrt{7.844^2 + 10.3294^2 - 2(7.844)(10.3294)\cos(1.28891^0)}$$

so, $\Delta V_1 = 2.4936 \text{ km/sec} = 8,977.085 \text{ km/hr}$

To find impulse needed to correct inclination at the apogee and transfer from Hohmann to the GEO orbit for an angle $(\beta - \alpha)$:

$$\Delta V_2 = \sqrt{V_{geo}^2 + V_a^2 - 2V_{geo}V_a\cos(\beta - \alpha)} = \sqrt{3.072^2 + 1.584^2 - 2(3.072)(1.584)\cos(13.711090)}$$

so, $\Delta V_2 = 1.578 \text{ km/sec} = 5682.388 \text{ km/hr}$

So, for this scenario, total ΔV is given by

$$\Delta V_1 + \Delta V_2 = 2.4936 + 1.578 = 4.0716 \text{ km/sec} = 14657.76 \text{ km/hr}$$

 $^2 \mathrm{See}$ end of this report for the MATLAB code.

4.2.4 Summary of scenarios to correct plane inclination

Comparing the ΔV from the above 3 scenarios, we see this:

scenario 1: 6.0218 km/sec scenario 2: 4.77575 km/sec scenario 3: 4.0716 km/sec

So, we can see that splitting the plane correction between the perigee and the apogee points leads to a more economical maneuver.

Hence Choose scenario 3 for the next sequence.

4.3 Calculate time to move from LEO to GEO

The vehicle was at the lines of node at t = 0, hence the time to reach GEO orbit is half the period of the Hohmann transfer orbit.

$$t = \frac{T_h}{2} = 18,916.77 \operatorname{sec} = 5.25 \operatorname{hr}$$

The above time is the same regardless if we inject at one end of the lines of nodes or at the other end. Also, this time is independent of what inclination the Hohmann transfer orbit was at the time of injection.

4.4 Decide when to inject to GEO. Calculate lead angle β and rendezvous with first satellite

In the previous step, the time it takes to move from LEO to GEO over a Hohmann orbit was found.

Now, find when to make this transfer. That is, we need to find the time to inject into the transfer orbit such that the overall ΔV is minimized.

The injection must occur when the space vehicle is on the line of nodes. Since this line is where the LEO and the desired GEO plane intersects at.

Hence, there are only 2 points on the LEO orbit that we can use to launch to GEO. (Both ends of the lines of nodes, at both sides of earth). Let me call one end of the lines of nodes, the *top* end, and the other end, the *bottom* end. Where the *top* end is that end which the spacecraft was at when it first reached LEO, i.e. at time=0.

In addition to the above restriction, if we want to eliminate the need to make any phasing loops when we arrive at the desired GEO orbit in order to rendezvous with the first target, then injection must occur only when the correct lead angle β is encountered. This additional synchronization requirement will turn out to be costly in time to achieve. If LEO and GEO orbits had been coplanar, then we can inject from any point on the LEO orbit as long as the lead angle β requirement is met. There will not be an additional requirement of the injection having to be from only two points in the LEO orbit.

We know that the synodic period between LEO and GEO is $\frac{T_{leo}T_{geo}}{T_{geo}-T_{leo}} = \frac{5189 \times 86391}{86391-5189} = 5520$ seconds, or 92 minutes. This means the LEO and the GEO objects will be aligned along a radial vector originating from the center of the earth once every 92 minutes.

But due to the restriction that this radial vector be only the lines of nodes of the space vehicle, using this synodic period is not too much help for me here.

So, what options do we have?

These are the options to investigate:

- 1. Inject from the top end of the lines of nodes. Reach GEO orbit and then phase-wait in that orbit to rendezvous with the target.
- 2. phase-wait in the LEO orbit until the correct lead angle β with the target is reached. Calculate this for when the spacecraft is on the top end of the lines of nodes.
- 3. The same as above, but for the case when the spacecraft is on the bottom end of the lines of nodes.
- 4. phase-wait in the LEO orbit until the lead angle is closest to β the first time this happens. (i.e. within the first 2π)

At the end, select the option which gives the smallest ΔV as long as the time cost is reasonable.

4.4.1 First option. Inject to GEO at t = 0

we know that at t = 0 we have this state as shown in figure below

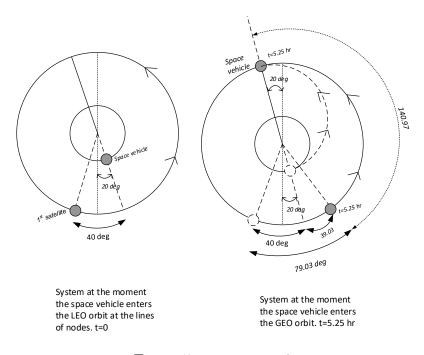


Figure 5: state at t = 0

From previous calculations, we found the time needed for space vehicle to reach GEO orbit is 18,916.77 sec. Hence, angle that the GEO satellite will travel in this time is found from

$$\frac{18916}{(0.9972696)(24)(60)(60)} = \frac{x}{2\pi}$$
$$x = 1.3793765886 \text{ rad}$$
$$= 79.03^{-0}$$

Hence When Vehicle reaches GEO, the first satellite will be (see diagram)

$$180 - (79.03 - 40) = 140.9675^{\circ}$$

behind the space vehicle.

Hence

$$\Delta L = -140.9675^{\circ}$$

The reason a minus sign is used, is because ΔL is measured positive if the change in longitude desired is eastwards, and since our target is behind us (westwards relative to the spacecraft), this change is negative.

$$\begin{split} -140.9675^0 &= -140.9675 \left(\frac{\pi}{180}\right) = -2.46 \text{ radians.} \\ \text{For } n &= 1, \ P_{ph} = \frac{-2.46}{1 \times \ 6.3} + 0.9972696 = 0.60676 \text{ days.} \\ \text{Hence } \overset{\circ}{L} &= \frac{\Delta L}{n \ P_{ph}} = \frac{-2.46}{1(0.60676)} = -4.054 \text{ radians/day} = -4.054 \left(\frac{180}{\pi}\right) = -232.2771 \\ \text{deg/day.} \\ \text{Hence } \Delta V &= 5.8 (-232.2771) = -1347.20718 \text{ m/sec} = -1.347 \text{ km/sec.} \\ \text{For } n &= 2, \ P_{ph} = \frac{-2.46}{2 \times \ 6.3} + 0.9972696 = 0.802 \text{ days.} \\ \overset{\circ}{L} &= \frac{-2.46}{2(0.802)} = -1.5337 \text{ rad/day or } -87.875 \text{ deg/day} \end{split}$$

Hence $\Delta V = 5.8(-87.875) = -509.675$ m/sec = -0.5 km/sec.

Continuing the above process, I obtain this table.

n	$\Delta V ({ m Km/s})$	Period of phase orbit (in days)	Total time in phasing period in days
1	-1.347	0.6067934	0.6067934
2	-0.509	0.80203	1.60406
3	-0.314	0.86711	2.60133
4	-0.2272	0.8996	3.5984
5	-0.17788	0.91917436	4.5958718
6	-0.1462	0.93219	5.59314
12	-0.07062	0.96472	11.57664
24	-0.0347	0.980999	23.543976
96	-0.0085	0.993202	95.347392

Any one of the above choices for n will achieve rendezvous with the first satellite.

Plotting ΔV against the total time in the phasing orbit results in the plot shown in figure below

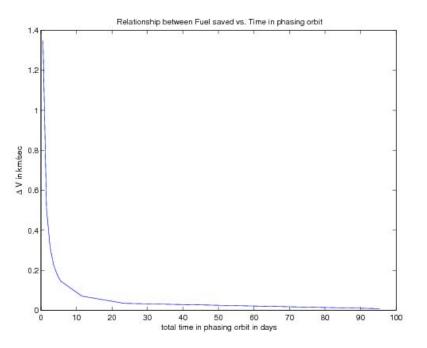


Figure 6: ΔV against the total time

Figure above shows that most saving in fuel is made by staying in the phasing orbit for less than 20 days. For more than 20 days, the additional saving in fuel is not justified by the time wasted in the phasing orbit.

Zooming in the region of interest in the plot shown in above figure results in figure below

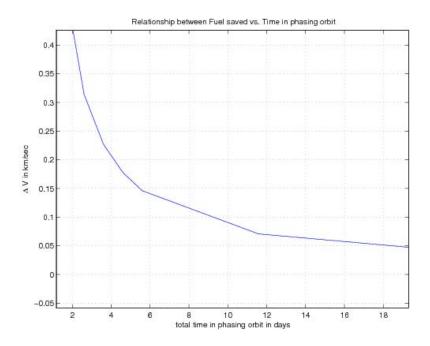


Figure 7: Zooming in the region of interest

From the above, it is clear that the slope after 6 days in the phasing orbit is less steep than earlier. The two options I see is to choose n = 6 and save some fuel, or choose n = 1 (the smallest possible value) and save time.

For n = 6

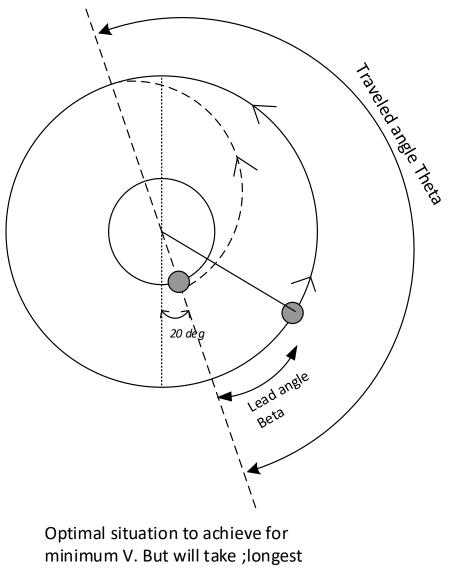
time = 5.59314 days
$$\Delta V = -0.1462$$
 km/sec.

For n = 1

time = 0.6067934 days
$$\Delta V = -1.347 \text{ km/sec}$$

Since the fuel saved is so small compared to the initial fuel needed to send the vehicle into GEO orbit, I decided to use the smaller time here.

4.4.2 second option: phase-wait in the LEO orbit until the correct lead angle β reached at one end of lines of nodes.



time to achieve.

Figure 8: second option

Looking at above figure, the problem can be seen as the following: we need to find the time it takes for the spacecraft to be at the top end of the lines of nodes when the target is at the correct lead angle β . Because in this case, the spacecraft can injects into the Hohmann orbit and will meet the target at the apogee. This would result in the

spacecraft not having to do any phase-waiting loops in the GEO orbit. Hence saving ΔV . A trade is made between time and ΔV .

How do we find this time value?

First, find β . To do this, equate travel time for target and spacecraft.

Travel time for spacecraft is half the Hohmann orbit period = 37,833 sec (see common calculations section for derivation)

Travel time x for target is found from

$$\frac{\text{distance travelled in radians}}{2\pi} = \frac{x}{\frac{\text{Period}}{x}}$$
$$x = \frac{P(\theta - \beta)}{2\pi}$$

Where period

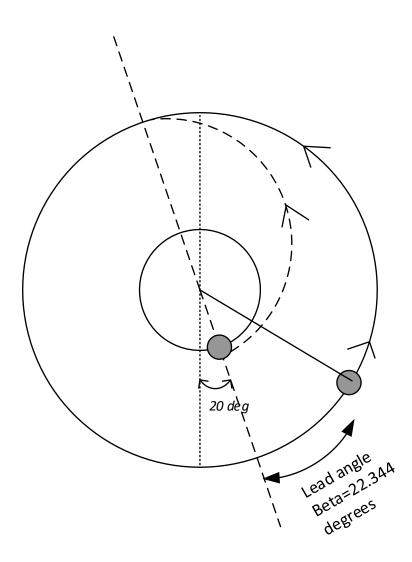
$$P = 2\pi \sqrt{\frac{r_{geo}^3}{\mu}}$$

= $2\pi \sqrt{\frac{42,238.145^3}{3.986012 \times 10^5}}$
= 86390 sec

Hence, and since $\theta = \pi$, when equating times of travel, we get this relation $\frac{86390(\pi-\beta)}{2\pi} = 37833$

Solve for β

$$37833(2\pi) = 86390(\pi - \beta)$$
$$\frac{37833(2\pi)}{86390} = \pi - \beta$$
$$\beta = \pi - \frac{37833(2\pi)}{86390}$$
$$= 0.389980 \text{ rad}$$
$$= 22.344^{0}$$



Optimal situation to achieve for minimum V.

Figure 9: Solve for β

So, now that we know β , we need to find when will the spacecraft be at the top end of the lines of nodes when the target has this β with it.

The angle that the target will move by for each one full period that the spacecraft makes in the LEO orbit is $T_{leo}\varpi$, where T_{leo} is the LEO period found in the common calculations section to be 5, 189 seconds, and ϖ is the average angular speed of the GEO target in radians per second which is 2π each 24 hrs (since GEO).

Exactly, $\varpi = \frac{2\pi}{0.9972696*24*60*60} = 0.00007292115609102$ rad per second.

Hence, target will travel 5, $189 \times 0.00007292115609102 = 0.378388$ rad = 21.68° for each one full LEO period.

Now that we know the angle the target will travel for each one full LEO period, we need to find how many times we have to do this so that target will end at the correct β location starting from t=0.

This becomes a simple counting problem. If we imagine a straight line, starting at t=0, and then we move a stick from its left end to its right end at an equal increments of 21.68 units, we just need to find when this stick will land at the correct point (or close enough) on the line where the point β is located.

When the stick reaches the right end of the line, we carry the reminder back to the start of the line and continue the process.

Figure below shows how to do this counting. Notice that the point I am interested in finding, which is the angle β , needs to be compensated for by adding the initial 40⁰ to it (since the counting is starting from the epoch). In other words, this is the degrees the spacecraft was ahead of the target when counting starts.

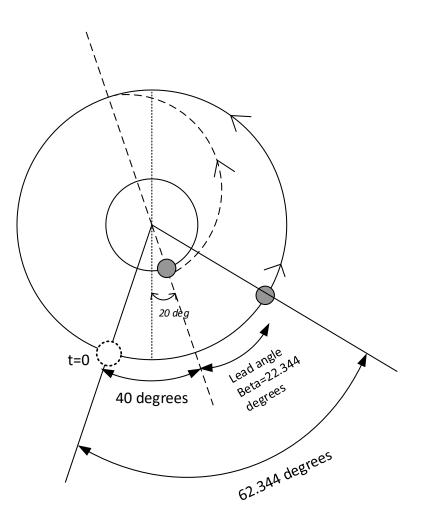


Figure 10: how to do this counting

So, the point on the line we are looking for is $40 + \beta = 40 + 22.344 = 60.344^{\circ}$, measured from t=0.

Figure below shows how the counting is actually done.

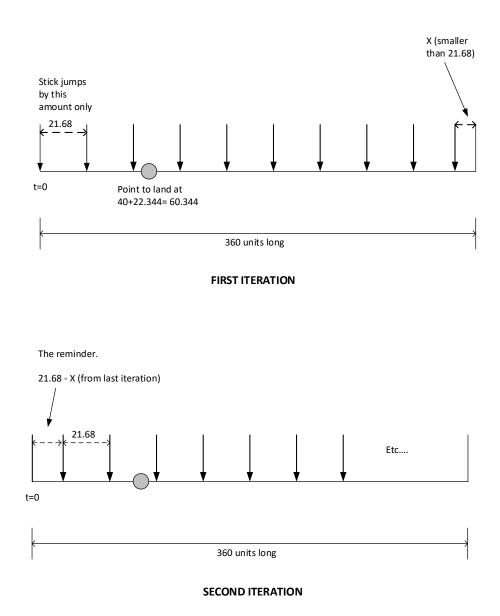


Figure 11: how the counting is actually done

I wrote a small MATLAB³ function called nma_findPointOnLine.m to do the counting. This function accepts as input the step size, the length of the line, and the distance we are looking for (60.344 in the above example), and how close to the target we want to be (the tolerance).

It returns the number of steps needed to achieve the synchronization.

This is an example call

³see appendix for source code

This is the result of the MATLAB function

tolerance (in degrees)	number of LEO periods needed	time to achieve (seconds)	time in hrs
1	36	$5189 \times 36 = 186804$	51.89
0.1	3274	$5189 \times 3274 = 16988786$	4719

It is clear that to achieve synchronization to 0.1 degree is too costly in time.

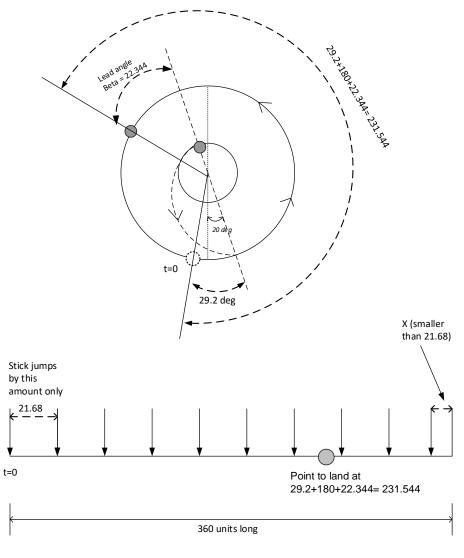
For the case of 1 degree tolerance, it will take 36 LEO loops to achieve the optimal situation with the target at the correct β .

This means, if we spend this time in LEO, we can inject and will meet the target at the same time when we reach GEO. Hence no additional ΔV would be needed in GEO to phase-wait. We have traded time for fuel.

4.4.3 Third option: phase-wait in the LEO orbit until the correct lead angle β reached at the other end of lines of nodes.

This case is the same as above, except now we want the spacecraft to be at the other end of the lines of nodes at injection. The only difference is that now t=0 have been shifted to become time after making one half LEO period. We can find this shift since we know the angle the target will travel in one half LEO period. We found from above that target will travel 21.68⁰ for one full LEO period, hence it will travel 10.84⁰ per half that period.

So, the only thing we need to do is determine where the point is located that we need to synchronize with, as illustrated by the figure below.



FIRST ITERATION

Figure 12: illustration of the above

So, use the same MATLAB function to find the number of full LEO rotations needed.

Hence, we see that for a tolerance of 1 degree, we have to wait 293 full LEO loops.

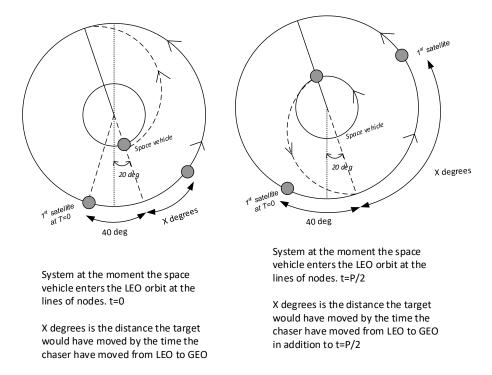
Compare this with the earlier case where we looked for the other end of the lines of nodes, which achieved the same synchronization for only 36 LEO loops. Hence this maneuver will not be accepted.

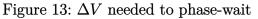
4.4.4 Fourth option. phase-wait in the LEO orbit until smallest difference to β reached first time.

To solve this problem, I will calculate the lead angle β with the first target for a number of time increments of 0.5*P* each, where *P* is the LEO period, and for each such time increment, will calculate where the first target will be at the end of a Hohmann transfer. Then will calculate the ΔV needed to phase-wait in GEO to close this final angle gap.

At
$$t = 0.5T_{leo}$$

Figure below illustrates this case.





Total time from epoch to reach $\text{GEO} = 0.5T_{leo} + 0.5T_h = 0.5(5189 + 37833) = 21,511 \text{ sec}$ Hence, angle that the GEO satellite will travel in this time is found from

$$\frac{21511}{(0.9972696)(24)(60)(60)} = \frac{x}{2\pi}$$
$$x = 1.568606 \text{ rad}$$
$$= 89.874^{-0}.$$

Hence When Vehicle reaches GEO, the first satellite will be

$$89.874 - 40 = 49.874^{\circ}$$

ahead.

At $t = T_{leo}$

Total time from epoch to reach $\text{GEO} = T_{leo} + 0.5T_h = 5189 + 0.5(37833) = 24,105.5 \text{ sec}$ Hence, angle that the GEO satellite will travel in this time is found from

$$\frac{24105.5}{(0.9972696)(24)(60)(60)} = \frac{x}{2\pi}$$
$$x = 1.7578 \text{ rad}$$
$$= 100.71^{0}$$

Hence When Vehicle reaches GEO, the first satellite will be

 $180 - (100.71 - 40) = 119.29^{\circ}$

behind

At $t = 1.5T_{leo}$

Total time from epoch to reach $\text{GEO} = 1.5T_{leo} + 0.5T_h = 1.5(5189) + 0.5(37833) = 26,700 \text{ sec}$

Hence, angle that the GEO satellite will travel in this time is found from

$$\frac{26700}{0.9972696 * 24 * 60 * 60} = \frac{x}{2\pi}$$
$$x = 1.946995 \text{ rad}$$
$$= 111.5546^{-0}$$

Hence When Vehicle reaches GEO, the first satellite will be

$$111.555 - 40 = 71.555^{\circ}$$

ahead.

Continue this process. The result is illustrated in this table

time	total time to reach GEO (sec)	ϕ target travelled angle	β lead angle	β sense relative to sp
0	18,916	79.03 ⁰	140.9675^{0}	behind
$0.5T_{leo}$	21, 511	89.874 ⁰	49.874^{0}	ahead
T_{leo}	24,105	100.71^{0}	119.29^{0}	behind
$1.5T_{leo}$	26,700	111.5546^{-0}	71.555^{0}	ahead
$2T_{leo}$	29, 294	122.3946^{0}	97.605^{0}	behind
$2.5T_{leo}$	31,889	133.2346^{0}	93.235^{0}	ahead
$3T_{leo}$	34,483	144.0746^{0}	75.9254^{0}	behind
$3.5T_{leo}$	37,078	154.9146^{0}	114.915^{0}	ahead
$4T_{leo}$	39,672	165.7547^{0}	54.2453^{0}	behind
$4.5T_{leo}$	42,267	176.5947^{0}	136.595^{0}	ahead
$5T_{leo}$	44,861	187.4347^{0}	32.5653^0	behind
$5.5T_{leo}$	47,456	198.2747^{0}	158.275^{0}	ahead
$6T_{leo}$	50,050	209.1147^{0}	10.8853^{0}	behind
$6.5T_{leo}$	52,645	219.9547^{0}	179.957^{0}	ahead
$7T_{leo}$	55, 239	230.7947^{0}	-10.7947^{0}	behind

From the above table, we see that the closest the target gets to the lines of nodes (within the first 2π) at the same time as the vehicle is 10.8853⁰ and is achieved after 6 periods in the LEO orbit.

What is left to do is to determine is the ΔV needed in the GEO orbit to wait-phase so as to close this final remaining $\Delta L = -10.8853^0$

Now, for the -10.8853° case calculate the cost ΔV for phase-waiting in the GEO orbit:

Let P_{ph} be the period of the phasing orbit while in the GEO orbit.

Let ω_E be the angular rate of axial rotation of the earth, which is 360.985647 deg/day = 360.985647 $\frac{\pi}{180} = 6.3$ radians/day.

Let P_o be the period of the geosynchronous orbit, which is 1436.068 min or 0.9972696 days.

Let ΔL be the change in longitude desired. Which we found it to be $-10^0 = -10(\frac{\pi}{180}) = -0.174533$ radians.

Let n be the number of revolutions spent in the phasing orbit.

Let \check{L} be the drift rate, positive eastwards.

First, find the period of the phasing orbit P_{ph} using the equation

$$P_{ph} = \frac{\Delta L}{n \; \omega_E} + P_o$$

for a specific *n*. Then solve for $\overset{\circ}{L}$ from the equation

$$\Delta L = \overset{\circ}{L}n \ P_{ph}$$

Then find ΔV corresponding to this $\overset{\circ}{L}$ from figure 7.14, Orbital Mechanics book or by using the relation

$$\Delta V = 5.8 \mathring{L}.$$

Where 5.8 is the slope of the line relating $\overset{\circ}{L}$ to ΔV . (The above slope is not exact, but it is close enough).

Try the above for a number of different values for n.

For
$$n = 1$$
, $P_{ph} = \frac{-0.174533}{1 \times 6.3} + 0.9972696 = 0.96957$ days.
Hence $\overset{\circ}{L} = \frac{\Delta L}{n P_{ph}} = \frac{-0.174533}{1(0.96957)} = -0.18$ radians/day $= -0.18 \left(\frac{180}{\pi}\right) = -10.31385$ deg/day.
Hence

 $\Delta V = 5.8(-10.31385) = -59.82033$ m/sec = -0.0598 km/sec.

I do not need to look for n=2 and higher for this case, since the saving in ΔV is clearly not worth spending an extra day for each additional increment in n. We see that the ΔV is very small already (0.0598 km/sec), and for larger n, it will only get smaller and smaller.

4.4.5 Summary of lead angle β scenarios

For first option, move to GEO without phase-waiting in LEO, and instead phase-wait in GEO, results in $\Delta V = -1.347$ with time spend (0.6067934)(24)(60)(60) = 52426 seconds, or 14 hrs and 33 minutes.

For the second and third options, the time spend is all in LEO, and zero ΔV was needed to phase-wait in GEO.

For 4th option, part of time spend is in LEO, and some part of time spend is in GEO. ΔV is not zero, but smaller than first option.

Notice that in this table, the ΔV cost refer only to the cost of phase wait in GEO to rendezvous with the first target.

scenario	$\Delta V \cos t$	LEO loop	total time cost to rendezvous
first. Inject at t=0	1.347 km/sec	0	52,426 (sec) = 14 hrs 33 minutes
second. optimal β at top end	0 km/sec	36	186,804 (sec) = 51.89 hrs = 2 days 3 hrs 53 n
third. optimal β at bottom end	0 km/sec	293	1,520,377 (sec)
fourth. smallest β at apogee	0.0598 km/sec	6	133,820 (sec) = 1 day, 13 hrs, 10 minutes, 20

Clearly option 3 is bad. With option 2, we get the same saving for much less time.

Between options 1 and 4, I prefer option 4, since for the cost of about 1.5 days, we reduced ΔV from 1.347 to 0.0598 km/sec.

So, the final choice is between option 4 and option 2.

With option 2, we have zero ΔV but we have to spend about 15 more hours in LEO to save 0.0598 km/sec. Is this good or not?

Compared to the ΔV needed to inject from LEO to GEO which is 4.0716 km/sec, this amount is 1.5% of that. It takes about 5 hrs to go from LEO to GEO. So, should I spend about 3 times as many hours to save 1.5% as many in fuel? (This is interesting that spending more time flying ends up saving fuel! only in space this is possible).

Will not consider option 2 as the time needed is not worth the saving in fuel.

Option 4 is selected

4.5 Rendezvous with the second satellite

Simply perform an in-orbit repositioning using a phasing orbit to rendezvous with the second satellite. The second satellite is 50^0 ahead of the first satellite (this is given), hence $\Delta L = 50^0 = 0.87267$ rad. Apply the same process of in-orbit repositioning to decide on the procedure to select. This table was generated:

n	$\Delta V ({ m Km/s})$	Period of phase orbit (in days)	Total time in phasing period in days
1	0.255	1.13578	1.13578
2	0.13596	1.066529	2.133058
3	0.093	1.04344	3.13032
4	0.0703	1.0318993	4.1275972
5	0.0566	1.0249734	5.124867
6	0.04746	1.020356	6.122136
12	0.02396	1.008812	12.105744
24	0.012	1.00304122	24.0729
96	0.003	0.998712	95.87635

Plotting ΔV against the total time in the orbit results in

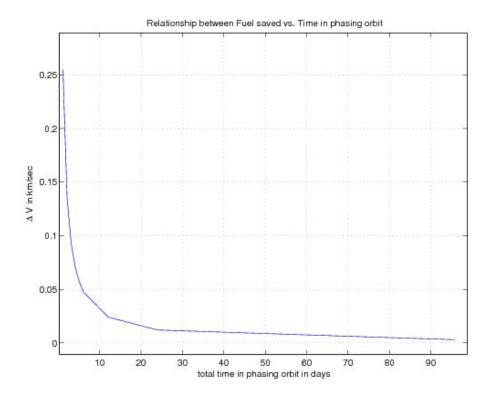


Figure 14: ΔV against the total time in the orbi

Similarly, at n = 6, the fuel saving is best for the time spend in the phasing orbit. This gives 6.122136 days in the phasing orbit, and

$$\Delta V = 0.04746 \text{ km/sec.}$$

However, since the fuel saved as a percentage of initial fuel is small, it does not seem that spending extra 6 days in orbit is worth it. I will use n = 1

Time = 1.3578 days
$$\Delta V = 0.255$$
 km/sec.

Also, since from the specifications, it seems that it says that the second satellite needs to be reached quickly, so I choose n = 1, the smallest possible value.

4.6 Stay locked in with the second satellite.

Stay in orbit for one complete orbit revolution. This adds one day to the total time in flight. No additional ΔV needed.

4.7 reposition to final destination.

Perform an in-orbit repositioning using a phasing orbit to reposition $\Delta L = 5^0 = 0.087266$ rad ahead.

Apply the same process of in-orbit repositioning to decide on the procedure to select. This table was generated:

n	$\Delta V ({ m Km/s})$	Period of phase orbit (in days)	Total time in phasing period in days
1	0.02868	1.0111213	1.0111213
2	0.0144	1.0041954	2.0083908
3	0.00965	1.00188684	3.00566052
4	0.007	1.00073253	4.00293012
5	0.0058	1.00003994	5.0001997
6	0.00484	0.9995782	5.9974692
12	0.00242	0.9984239	11.9810868
24	0.001211	0.99784675	23.948322
96	0.0003	0.9974138	95.7517248

As before, choose n = 1 with

Time = 1.0111213 day
$$\Delta V = 0.02868$$
 km/sec.

This adds little over one day to the total flight time.

This completes the required sequence.

5 Summary and Results

See figure below for the final decision tree.

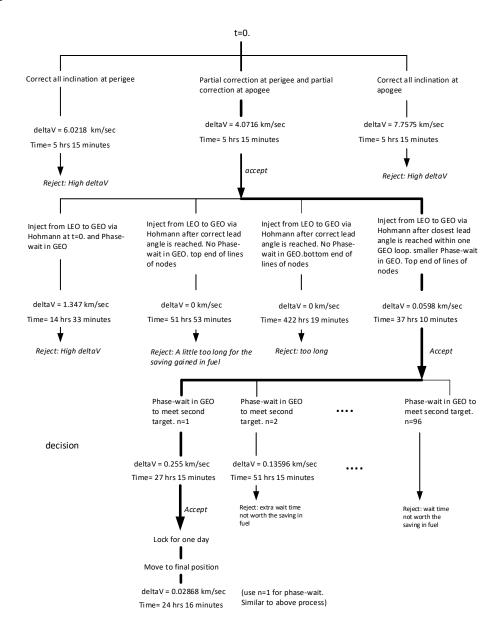


Figure 15: final decision tree

This is a summary, in table format, of the entire orbiting sequences using the most ΔV optimal maneuvers selected out of the different scenarios shown above. This table below is the final result of selecting the best option out of each phase. This table only shows the ΔV needed starting from LEO orbit. It does not include the ΔV needed to

reach LEO which I will add next.

sequence: Inject to LEO. wait in LEO for 6 full orbits. Perform Hohmann transfer with partial plane correction at each end. phase-wait in GEO to rendezvous with first target. Phase-wait in GEO to rendezvous with second target. Lock into second target. Reposition to final location in GEO.

sequence	$\Delta V (\rm km/sec)$	duration
wait in LEO	0	50,050 (sec)
Hohmann transfer	4.0716	18,916 (sec)
Rendezvous with 1st target	0.0598	83,770 (sec)
Rendezvous with 2nd target	0.255	98,131 (sec)
lock with 2nd target	0	86,400 (sec)
position to final destination	0.02868	87,360 (sec)
TOTAL	4.41508	424,627 (sec)=117.95 hrs = 4 days 21 hrs 57 min

To reach LEO, we have found that $V_{leo} = 7.844$ km/sec. To be more realistic, we need to account for the gravitational loss and drag. Typical time to reach LEO is about 2 minutes or 120 seconds. Hence additional ΔV to account for gravity loss is $g\Delta t = 9.8(120) = 1176$ m/sec = 1.176 km/sec.

For the drag effect, it of course depends on the rocket cross sectional area, the drag coefficient and air density. A typical value I have seen in the literature for rockets is to use 5% of the LEO velocity to account for drag. Hence an additional ΔV for drag will be 5% of 7.844 km/sec or 0.392 km/sec.

Hence total ΔV to reach LEO = 7.844 + 1.176 + 0.392 = 9.412 km/sec

Hence the total ΔV to achieve final position of spacecraft is

4.41508 + 9.412 = 13.827 km/sec

see figure below for a graphical display of the deltaV used at each stage.

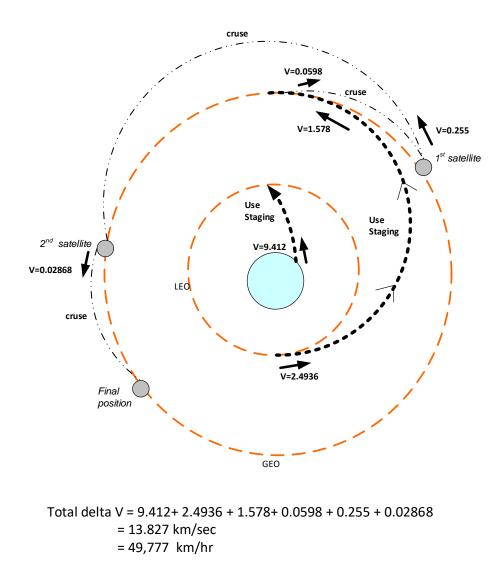


Figure 16: graphical display of the deltaV used at each stage

6 Appendix

This is the MATLAB function that solves for α to find what partial correction in inclination angle can be done at the perigee for the Hohmann transfer.

Caller script:

```
function nma_testfindAlphaForMinDeltaV
r0 = 6378.145;
r1 = 100+r0;
r2 = 35860+r0;
beta = 15;
```

```
mu = 3.986012*10<sup>5</sup>;
alpha = nma_findAlphaForMinDeltaV(r1,r2,beta,mu)
```

This is the function that solves for alpha.

```
function alpha=nma_findAlphaForMinDeltaV(r1,r2,beta,mu)
%function alpha=nma_findAlphaForMinDeltaV(r1,r2,beta,mu)
%
% Finds the minimum alpha (initial inclination correction)
% for an orbit relative to a larger circular orbit.
% see design note for more details.
%
%INPUT:
% r1: The radius of the smaller circular orbit
% r2: the radius of the larger circular orbit
  beta: the angle, in degrees, in which the two
%
%
         circular orbits are non co-planers to
%
        each others.
% mu: gravitational constant
%
% OUTPUT:
%
  alpha: The angle in degrees to use for initial
%
          correction such that minimum delta V is
%
          obtained to move from the smaller circular
%
          orbit to the larger circular orbit
%
% Author: Nasser Abbasi
%
          May 19,2003
a = (r1+r2)/2;
rp = r1;
ra = r2;
Va=sqrt( mu*(2/ra - 1/a ));
Vp=sqrt( mu*(2/rp - 1/a ));
Vc1=sqrt( mu/r1 );
Vc2=sqrt( mu/r2 );
beta=beta*pi/180;
\end{Verbatim}
\newpage\begin{Verbatim}
root(1)=0.1*beta;
```

```
keepLooking = true;
i=0;
while(keepLooking)
i=i+1;
root(i+1)=root(i)- ( F(Vc1,Vc2,Vp,Va,root(i),beta)/dF(Vc1,Vc2,Vp,Va,root(i),beta) );
root(i+1)
if( abs ( (root(i+1) - root(i)) / root(i+1) ) * 100 < 0.001 )</pre>
keepLooking=false;
else
if( ( root(i+1) * root(i) )<0.0)</pre>
error('jumped out of root');
end
end
if(i>100)
error('Failed to converge');
end
end
alpha=root(end);
alpha=alpha*180/pi;
%
%
function v=F(Vc1,Vc2,Vp,Va,alpha,beta)
v=Vc1*Vp*sin(alpha)*sqrt(Vc2^2+Va^2-2*Vc2*Va*cos(beta-alpha)) ...
- Vc2*Va*sin(beta-alpha)*sqrt(Vc1^2+Vp^2-2*Vc1*Vp*cos(alpha));
%
%
function v=dF(Vc1,Vc2,Vp,Va,alpha,beta)
v=Vc1*Vp*cos(alpha)*sqrt(Vc2^2+Va^2-2*Vc2*Va*cos(beta-alpha)) ...
+Vc1*Vp*sin(alpha)* ...
(-Vc2*Va*sin(beta-alpha)/sqrt(Vc2^2+Va^2-2*Vc2*Va*cos(beta-alpha))) ...
- (Vc2*Va*cos(beta-alpha)*sqrt(Vc1^2+Vp^2-2*Vc1*Vp*cos(alpha)) ...
+ Vc2*Va*sin(beta-alpha)*...
( Vc1*Vp*sin(alpha) / sqrt(Vc1^2+Vp^2-2*Vc1*Vp*cos(alpha))));
\end{Verbatim}
```

```
\newpage\begin{Verbatim}
```

```
function nSteps=nma_findPointOnLine(stepSize,lineLength,target,tol)
%function nSteps=nma_findPointOnLine(stepSize,lineLength,target,tol)
%
% Function to find how many steps needed to reach withing tolearance
% close to a point on a line by taking fixed number of steps. Line wrapes
% around.
%
%INPUT:
%
   stepSize: the step size
%
   lineLength: The line length
%
   target: The distance from leftend of line that we want to reach
%
   tol: tolerance in abs. value
%
%OUTPUT:
%
   nSteps: number of steps needed
%
%Author Nasser Abbasi. May 22, 2003
%
currentDist = 0;
nSteps
            = 0;
while true
currentDist = currentDist+stepSize;
nSteps = 1;
while(currentDist <= lineLength)</pre>
if( abs (currentDist-target) < tol )</pre>
return:
end
rem = lineLength-currentDist;
if(rem < stepSize)</pre>
currentDist = stepSize-rem;
else
currentDist = currentDist+stepSize;
end
nSteps = nSteps+1;
end
end
```